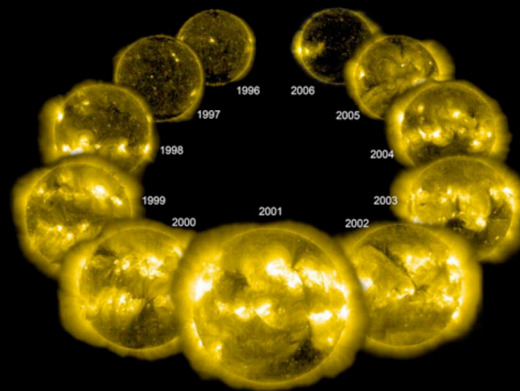
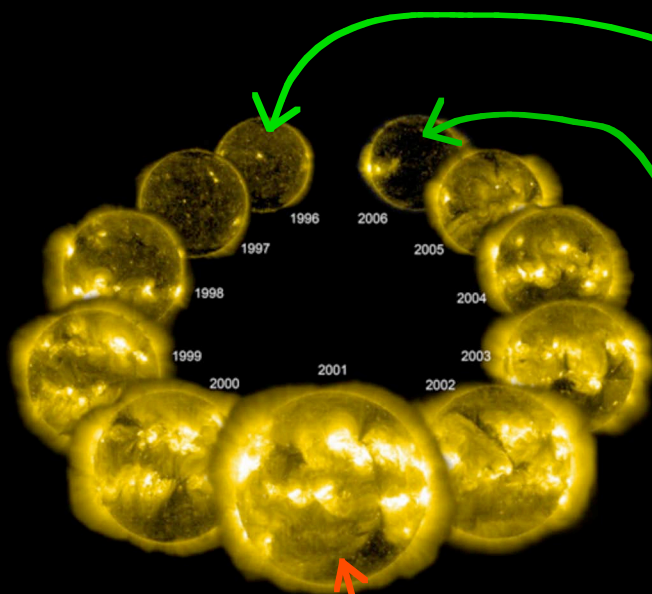


Dynamic Sun II

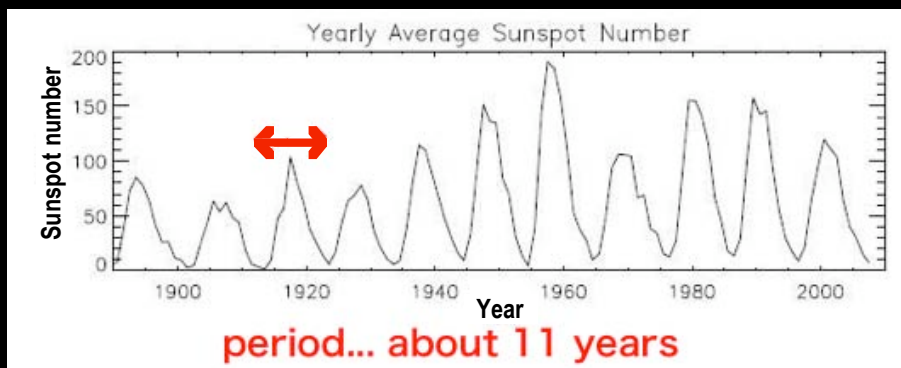
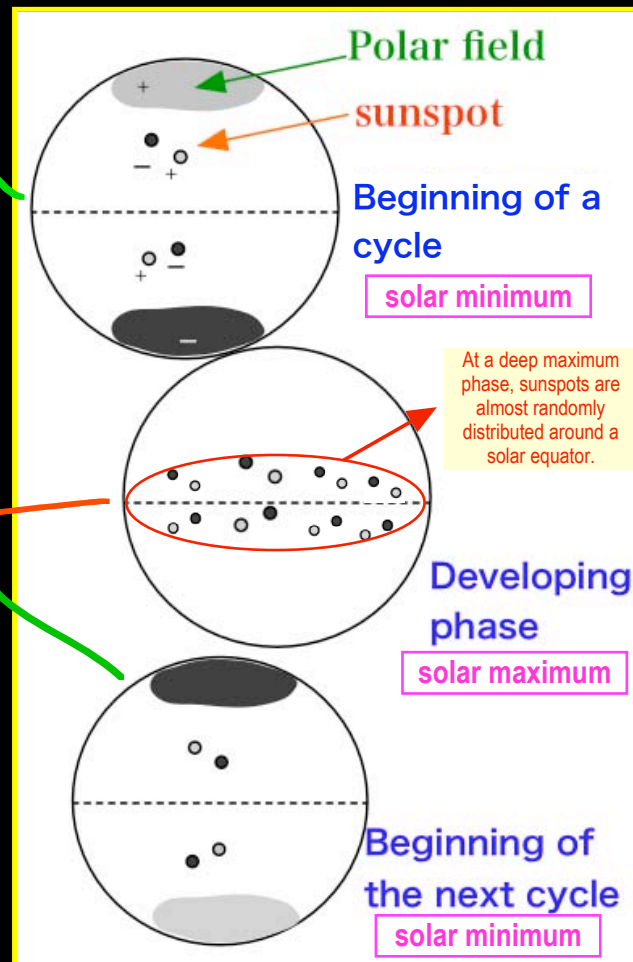
(long time-scale phenomena: ~ years)



Solar cycle: periodic variations of solar global activity

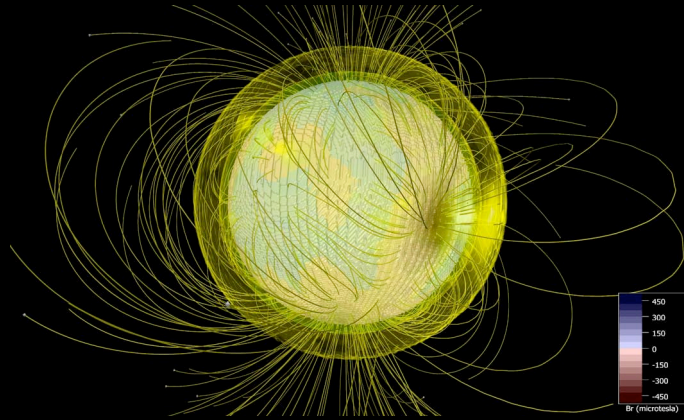


Hale's law on solar magnetic cycle



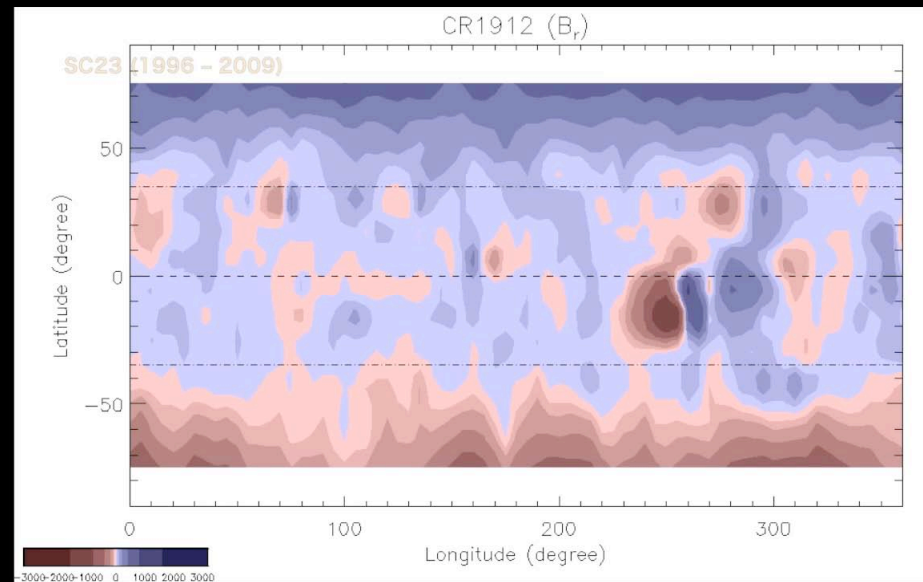
Variations of the Sun's large-scale magnetic field

<http://163.180.179.74/~magara/Download/SDMS-2021-02d.pdf>



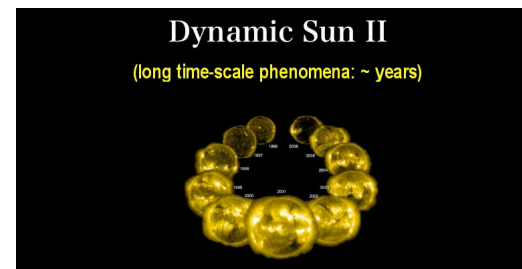
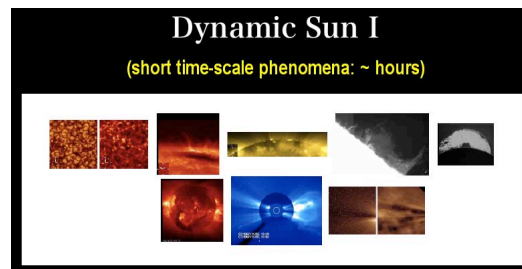
CR1912
(1996 07/25 - 08/21)

WSO synoptic maps



unit: μT

What produces solar dynamic phenomena?



Magnetic fields play important roles in producing solar dynamic phenomena.

- **Cyclic amplification of magnetic fields in the solar interior (*dynamo*)**
=> Produces long-term activity variations known as solar cycle
- **Transport of magnetic fields through the convection zone (*magnetoconvection*)**
=> Lifts magnetized plasmas against solar gravity via magnetic buoyancy
- **Emergence of magnetic fields into the solar atmosphere (*flux emergence*)**
=> Forms magnetic structure on the Sun (e.g. sunspot, prominence/filament, sigmoid)
- **Diffusion of magnetic fields in the solar atmosphere (*release of magnetic energy*)**
=> Produces explosive phenomena via magnetic reconnection (e.g. flare, jet, coronal heating?)
- **Ejection of magnetic fields into the interplanetary space (*removal of magnetic fields from the Sun*)**
=> Produces outflow/eruptive phenomena (e.g. solar wind, coronal mass ejection)