

## Summary of motions of a gyrating particle

**Standard form:**  $m \frac{d\mathbf{v}}{dt} = q \mathbf{v}_\perp \times \mathbf{B}_0 + \mathbf{F}_0$

$q \mathbf{v}_\perp \times \mathbf{B}_0 \longrightarrow$  gyration

$\mathbf{F}_0$  (external force)  $\longrightarrow$  drift, mirror effect

$q \mathbf{v}_\perp \times \mathbf{B}$  ( $B$  changes in  $\perp$ -direction)  $\longrightarrow$   $q \mathbf{v}_\perp \times \mathbf{B}_0 + \mathbf{F}_0$

$\mathbf{F}_0 = -\frac{m v_G^2}{2 B_0} \nabla_\perp B$

$q \mathbf{v}_\perp \times \mathbf{B}$  ( $B$  changes in  $\parallel$ -direction)  $\longrightarrow$   $q \mathbf{v}_\perp \times \mathbf{B}_0 + \mathbf{F}_0$

$\mathbf{F}_0 = -\frac{m v_G^2}{2 B_0} \nabla_\parallel B$

**Fundamental form:**  $\mathbf{F}_0 = -\nabla(-\boldsymbol{\mu} \cdot \mathbf{B})$   $|\boldsymbol{\mu}| \sim \text{const.}$

$\boldsymbol{\mu} = -\frac{m v_G^2 / 2}{B_0} \hat{\mathbf{b}}, \hat{\mathbf{b}} \equiv \frac{\mathbf{B}_0}{|\mathbf{B}_0|}$  (unit vector of  $\mathbf{B}_0$ )

diamagnetic

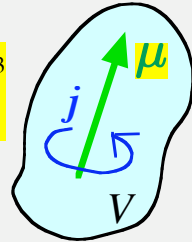
# Magnetic moment of a charge particle

# Magnetic (dipole) moment...

Definition of **magnetic moment**:

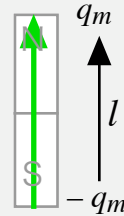
$$\boldsymbol{\mu} \equiv \frac{1}{2} \int_V \mathbf{r} \times \mathbf{j} \, d\tau^3$$

current-based definition



$$\boldsymbol{\mu} \equiv q_m \mathbf{l}$$

magnetic charge-based definition



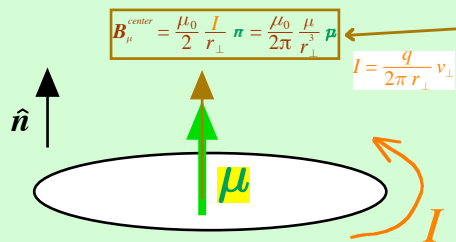
A diagram showing a magnetic dipole (red loops) in an external magnetic field  $\mathbf{B}_0$  (blue arrows). The dipole moment  $\boldsymbol{\mu}$  is shown pointing upwards. The magnetic field  $\mathbf{B}_\mu$  is also shown.

potential energy:  $U_M = -\boldsymbol{\mu} \cdot \mathbf{B}_0(\mathbf{r})$   
 force:  $\mathbf{F} = -\nabla U_M$   
 torque:  $\mathbf{N} = \boldsymbol{\mu} \times \mathbf{B}_0$

$B_r = \frac{\mu_0}{4\pi} \left[ \frac{3\mathbf{r}(\boldsymbol{\mu} \cdot \mathbf{r}) - r^2 \boldsymbol{\mu}}{r^3} + \frac{8\pi}{3} \boldsymbol{\mu} \delta^3(\mathbf{r}) \right]$

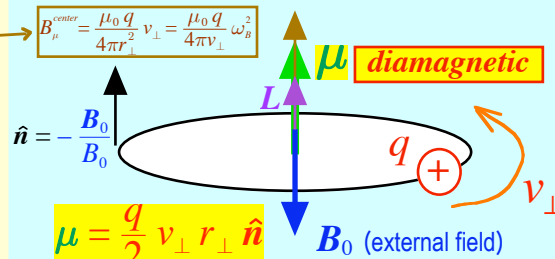
Magnetic moments caused by **rotating motions of charged particle**:

## Circular current (conduction current)



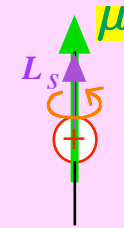
$$\boldsymbol{\mu} = I S \hat{n} = I \pi r_\perp^2 \hat{n}$$

## Gyrating charged particle



$$\begin{aligned} \boldsymbol{\mu} &= \frac{q}{2} v_\perp r_\perp \hat{n} \\ &= \frac{q}{2m} \mathbf{L} \quad v_\perp r_\perp \hat{n} = \mathbf{L} / m \\ &= I \pi r_\perp^2 \hat{n} \quad I = \frac{q}{2\pi r_\perp} v_\perp \\ &= -\frac{m v_\perp^2 / 2}{B_0} \frac{B_0}{B_0} \quad r_\perp = \frac{v_\perp}{\omega_B} \\ &= -\frac{1}{2\pi} \frac{q^2}{m} \Phi_B \frac{B_0}{B_0} \quad \Phi_B = \pi r_\perp^2 B_0 \end{aligned}$$

## Spinning charged particle



$$\boldsymbol{\mu} = g \frac{q}{2m} \mathbf{L}_S$$

spin angular momentum  $\mathbf{L}_S$   
 ... quantized by  $\hbar$  or  $\hbar/2$

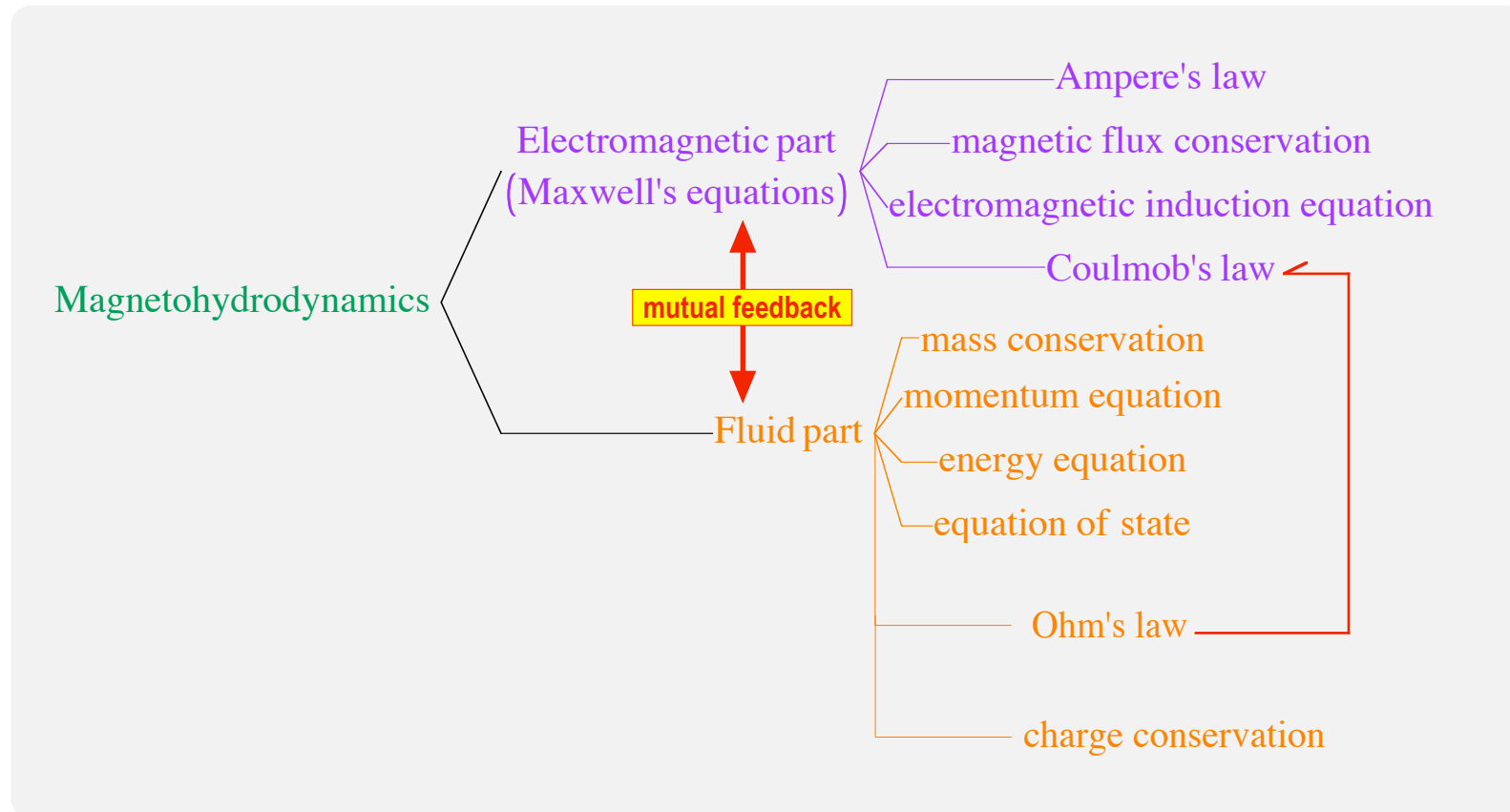
g... g-factor

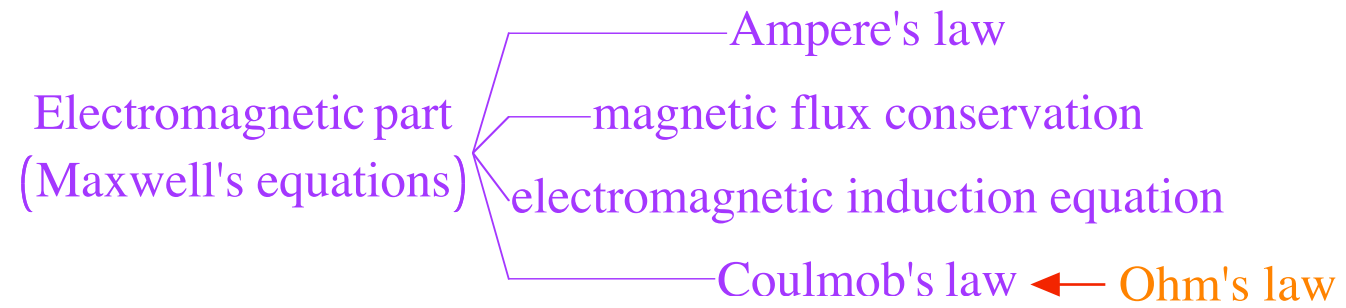
# Magnetohydrodynamics

A fluid theory of plasmas

## MHD equations...

Plasma (fluid) affects electromagnetic fields, while electromagnetic fields affect plasma.





**(Plasma => Fields)**

**Maxwell's equations** (determine electromagnetic fields from plasma quantities)

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{j} + \frac{1}{c^2} \frac{\partial \mathbf{E}}{\partial t} \quad \dots \text{Generalized Ampere's law}$$

MKS unit

$$\nabla \cdot \mathbf{B} = 0 \quad \dots \text{Conservation of magnetic flux}$$

$$\nabla \times \mathbf{E} = - \frac{\partial \mathbf{B}}{\partial t} \quad \dots \text{Electromagnetic induction equation (Faraday's law)}$$

$$\nabla \cdot \mathbf{E} = \frac{\rho_c}{\epsilon_0} \quad \dots \text{Coulomb's law}$$

$$\epsilon_0 = 8.854 \times 10^{-12} \text{ F m}^{-1} \quad \dots \text{permittivity of free space}$$

$$\mu_0 = 4\pi \times 10^{-7} \text{ H m}^{-1} \quad \dots \text{magnetic permeability}$$

$$c = \frac{1}{\sqrt{\epsilon_0 \mu_0}} = 2.998 \times 10^8 \text{ m s}^{-1} \quad \dots \text{speed of light}$$

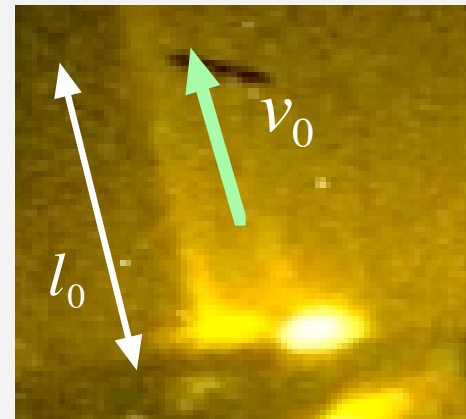
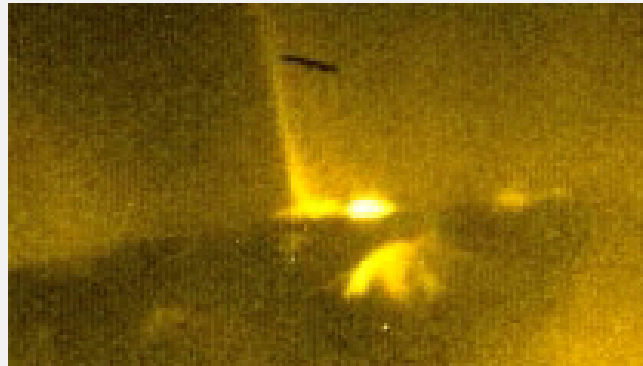
**Unit:**

$E$ ... V/m,  $B$ ... T (=  $10^4$  G),  $j$ ... A/m<sup>2</sup>

## Typical scales used in MHD

MHD phenomena... *large-scale, slowly evolving phenomena*

e.g. solar coronal jet



Typical scale:

length...  $l_0 \sim 10^4$  km

velocity...  $v_0 \sim 100$  km/s

time...  $t_0 = \frac{l_0}{v_0} \sim 100$  s

$r_g$  (gyration radius of a proton)  $\sim 1$  m  
 $\tau_g$  (gyration period of a proton)  $\sim 10^{-4}$  s  
in the solar corona