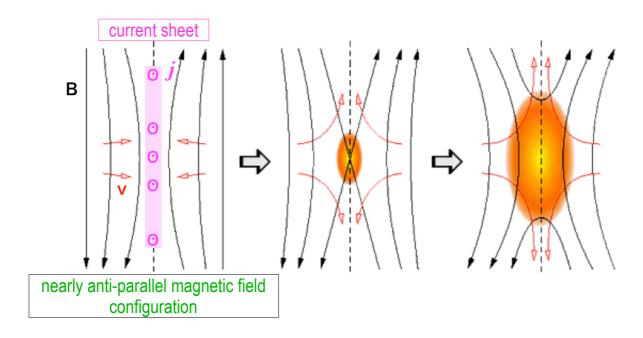
# Diffusion of magnetic fields in the solar atmosphere

# Magnetic reconnection

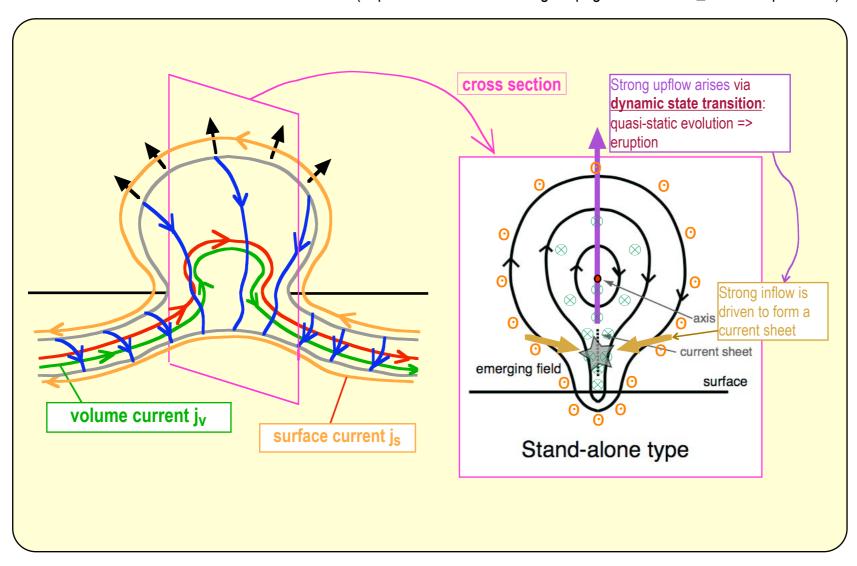
When a nearly anti-parallel magnetic field configuration exists, strong cross-field electric current flows there (current sheet).



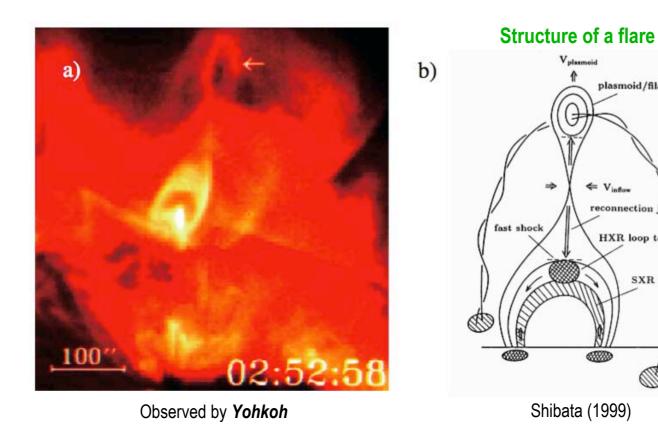
In the current sheet, energy conversion called 'magnetic reconnection' may proceed, which efficiently converts free magnetic energy to kinetic and thermal energy. It may also accelerate selected particles via electric field involved in the reconnection.

# How is a current sheet formed in an emerging twisted flux tube?

(http://163.180.179.74/~magara/page31/Research\_evolution-path.html)



# **Solar flare:** magnetic reconnection is expected to occur...



### **Properties of flares**

**Spatial size: 10^4 – 10^5 km** 

**Duration: 10^3 - 10^4 sec** 

Released energy: 10^28 – 10^32 erg/event  $\leftarrow$   $L_o \sim 4 \times 10^{33}$  erg/s

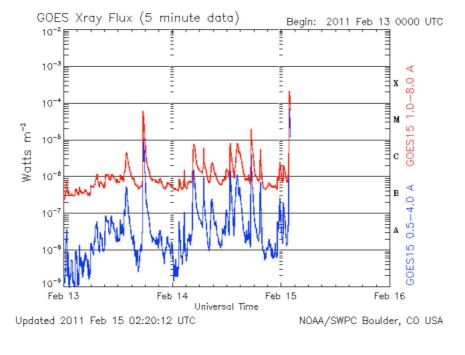
plasmoid/filament

reconnection jet

HXR loop top source

SXR loop

# Classification of flares based on X-ray emission



#### Peak value of X-Ray flux (1 – 8 Å, W/m<sup>2</sup>)

A 
$$10^{-8} - 10^{-7}$$

**B** 
$$10^{-7} - 10^{-6}$$

$$\mathbf{C} \quad 10^{-6} - 10^{-5}$$

$$M 10^{-5} - 10^{-4}$$

$$X > 10^{-4}$$

# Ejection of solar magnetic fields into the interplanetary space

# **Properties of CMEs**

- Occurrence rate: 1 (solar minimum) ~ 3 (maximum) /day
- Average velocity: 450 km/s

slow CME... high latitude, gradually accelerated fast CME... active region, ballistically accelerated

- Total mass loss: 10<sup>13</sup> 16 g/event ← solar-wind mass loss rate: 10<sup>12</sup> g/s
- Kinetic (mechanical) energy: 10^27 32 erg/event  $\leftarrow$   $L_o \sim 4 \times 10^{33}$  erg/s

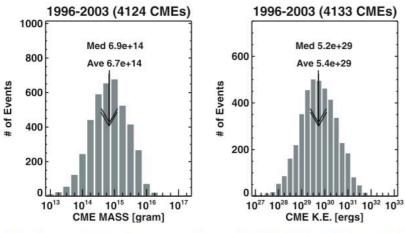


Table 1: Some Average Characteristics of Coronal Mass Ejections

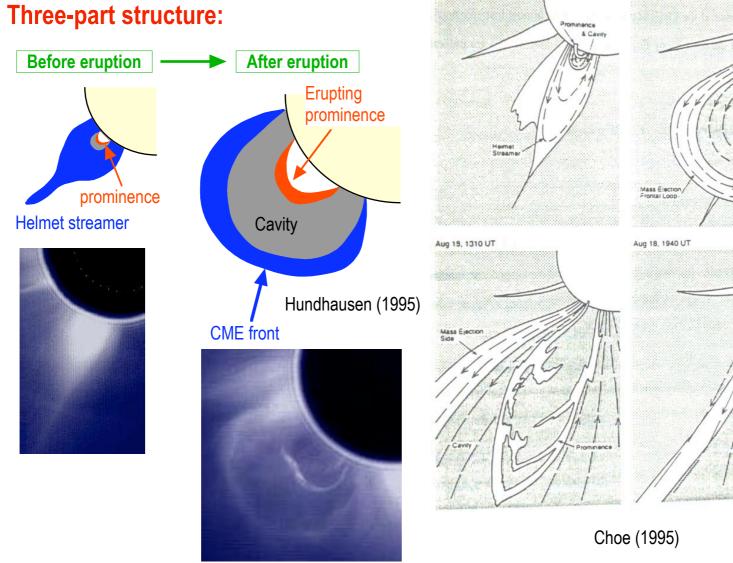
	Skylab (1973-74)	Solwind (1979-80 & 1984-85)	SMM (1980, 1984-89)
Angular Size	42°	43°	47°
Speed	$470 \text{ km sec}^{-1}$	$460 \text{ km sec}^{-1}$	350 km sec <sup>-1</sup>
Mass	WARRANT STORY (TO STORY TO	$4.0 \times 10^{15} \text{ gm}$	$3.3 \times 10^{15} \text{ gm}$
Kinetic Energy		$3.4 \times 10^{30} \text{ ergs}$	$6.7 \times 10^{30} \text{ ergs}$
Potential Energy			$7.1 \times 10^{30} \text{ ergs}$
Mechanical Energy			$1.38 \times 10^{31} \text{ ergs}$

Hundhausen (2000)

Figure 2. CME mass (left), and kinetic energy (right) of SOHO CMEs for the period 1996–2003.

Gopalswamy (2006)

### Structure of a CME



Aug 17, 2258 UT

Aug 18. 1215 UT

Expanded Prominence