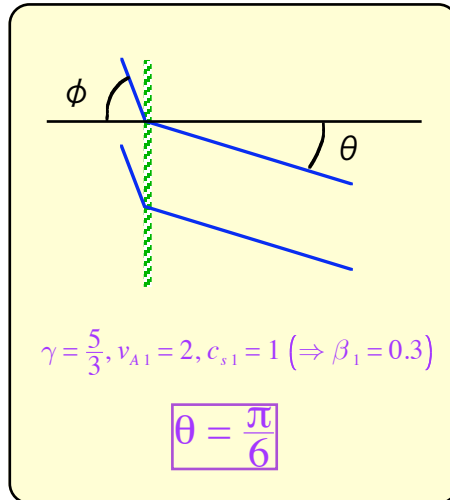
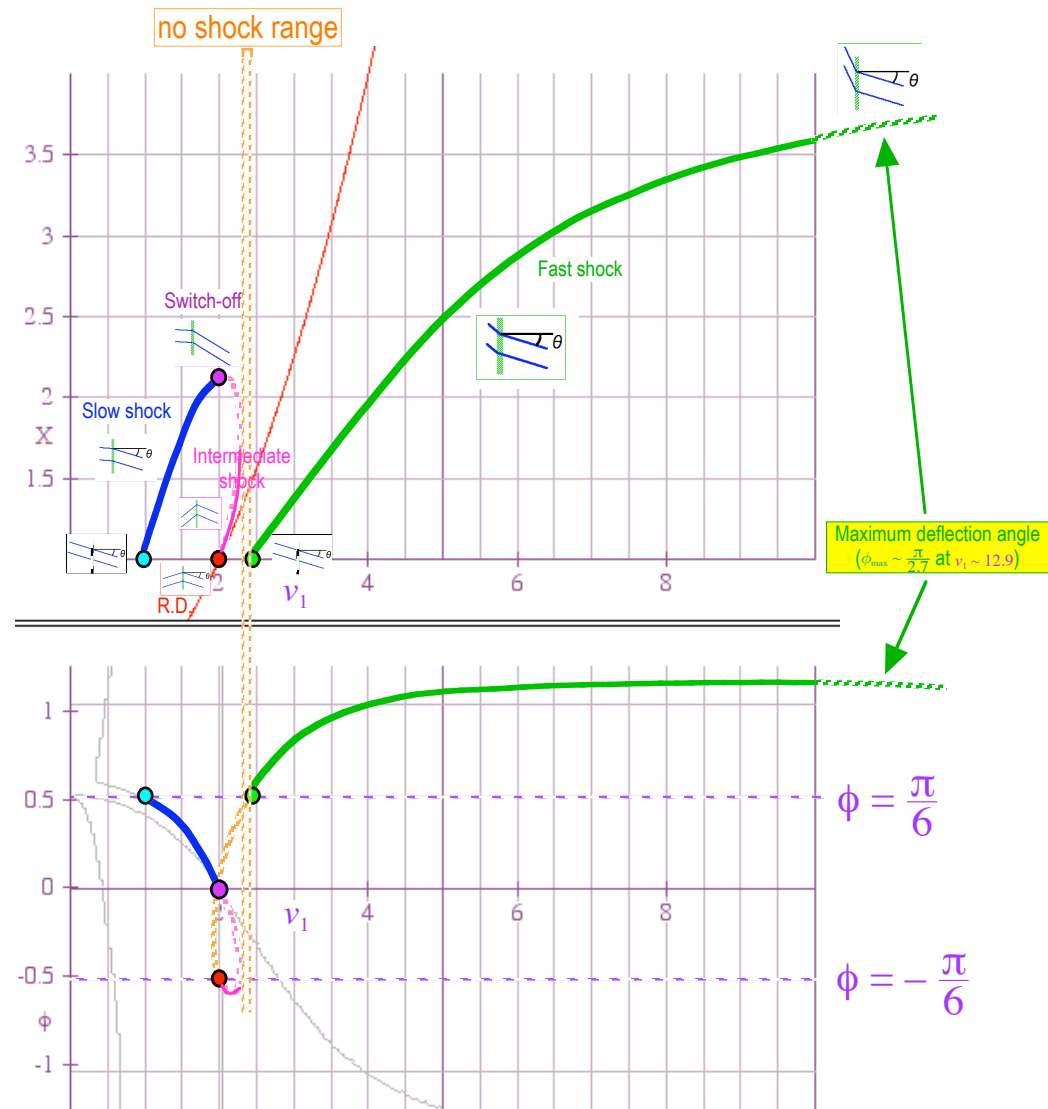


Angle relation (θ, ϕ)



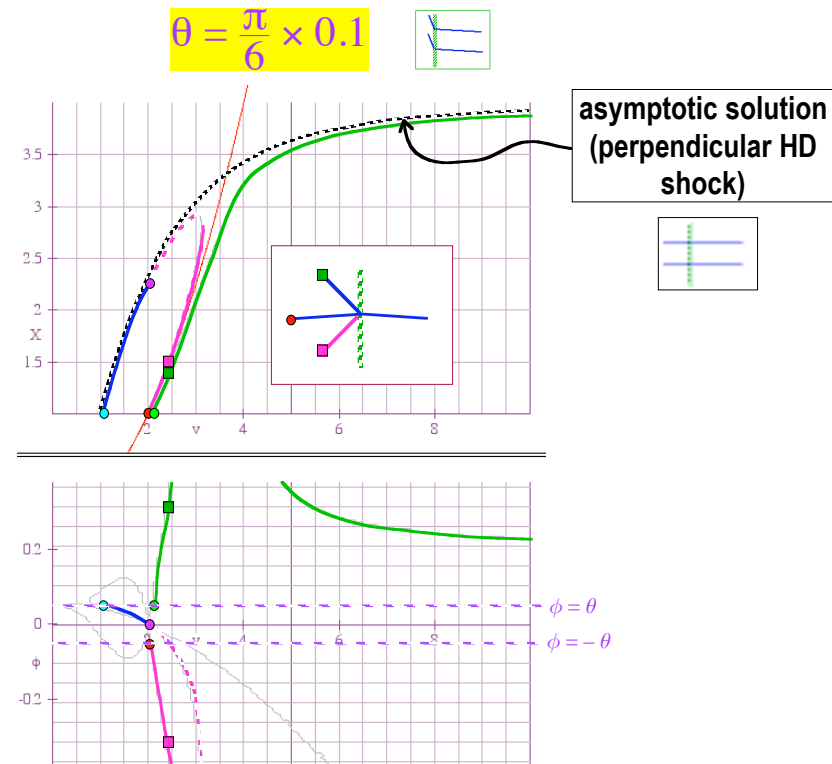
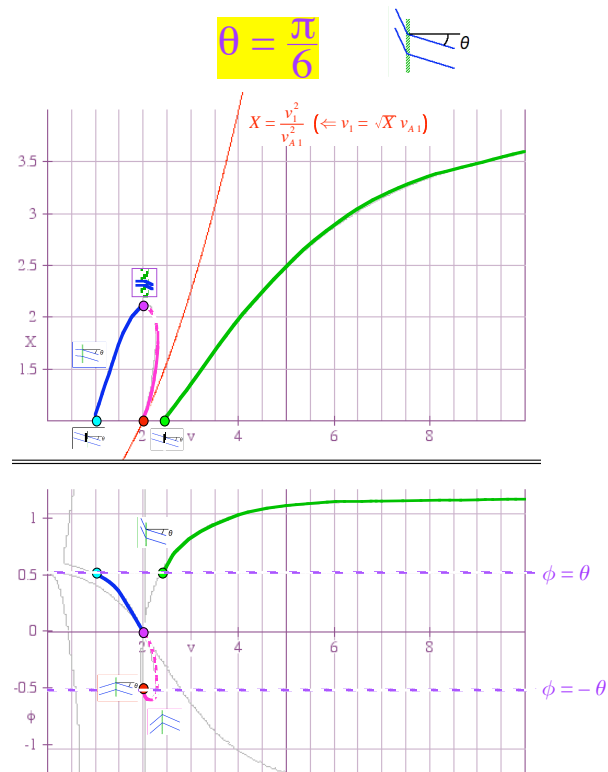
Fast shock $v_1 \geq \sqrt{X} v_{A1}$
 Slow shock $v_1 \leq v_{A1}$
 Rotational discontinuity $v_1 = v_{A1}$

Switch-off shock with $v_1 = v_{A1}$
 $(v_{x1} = \frac{B_{x1}}{\sqrt{\mu \rho_1}} = \frac{B_{x2}}{\sqrt{\mu \rho_1}})$



Transition from normal fast shock to switch-on shock

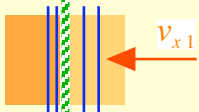
$$\gamma = \frac{5}{3}, \beta = 0.3, v_{A1} = 2, c_{s1} = 1$$



Several limiting cases of oblique MHD shock

Fast shock...

when $\theta = \frac{\pi}{2}$, perpendicular MHD shock

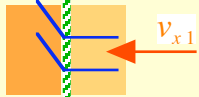


$$1 < X(v_1) < \frac{\gamma + 1}{\gamma - 1}$$

no upper limit of v_{x1}

when $\theta \rightarrow 0$ (low-beta case), switch-on shock

$$v_{A1} < v_{x1} \leq v_{A1} \sqrt{\frac{\gamma + 1 - 2c_s^2/v_{A1}^2}{\gamma - 1}}$$

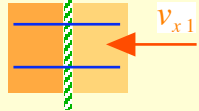


$$1 < X(v_1) \leq \frac{\gamma + 1 - 2c_s^2/v_{A1}^2}{\gamma - 1}$$

low-beta

when $\theta = 0$, perpendicular HD shock

In a low-beta case with $v_{x1} > \sqrt{X} v_{A1}$, it becomes perpendicular HD shock.
In a high-beta case with $v_{x1} > c_{s1} > v_{A1}$, it becomes perpendicular HD shock.



$$\frac{\gamma + 1 - 2c_s^2/v_{A1}^2}{\gamma - 1} < X(v_1) < \frac{\gamma + 1}{\gamma - 1}$$

low-beta

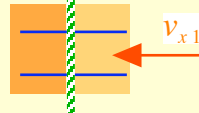
$$1 < X(v_1) < \frac{\gamma + 1}{\gamma - 1}$$

no upper limit of v_{x1}
high-beta

Slow shock...

when $\theta = 0$ (low-beta case), perpendicular HD shock

In a low-beta case with $v_{A1} > v_{x1} > c_{s1}$, it becomes perpendicular HD shock.




$$1 < X(v_1) \leq \frac{\gamma + 1}{\gamma - 1 + 2c_s^2/v_{A1}^2}$$

low-beta

Tangential discontinuity...

simple boundary between two regions where total pressure (magnetic pressure + gas pressure) is the same


$$p_1 + \frac{B_1^2}{2\mu} = p_2 + \frac{B_2^2}{2\mu}$$


X is variable

$v_{x1} = 0, v_{x2} = 0$

Contact discontinuity...

simple boundary between two regions where density and temperature take different values

$$p_1 = p_2$$


X is variable

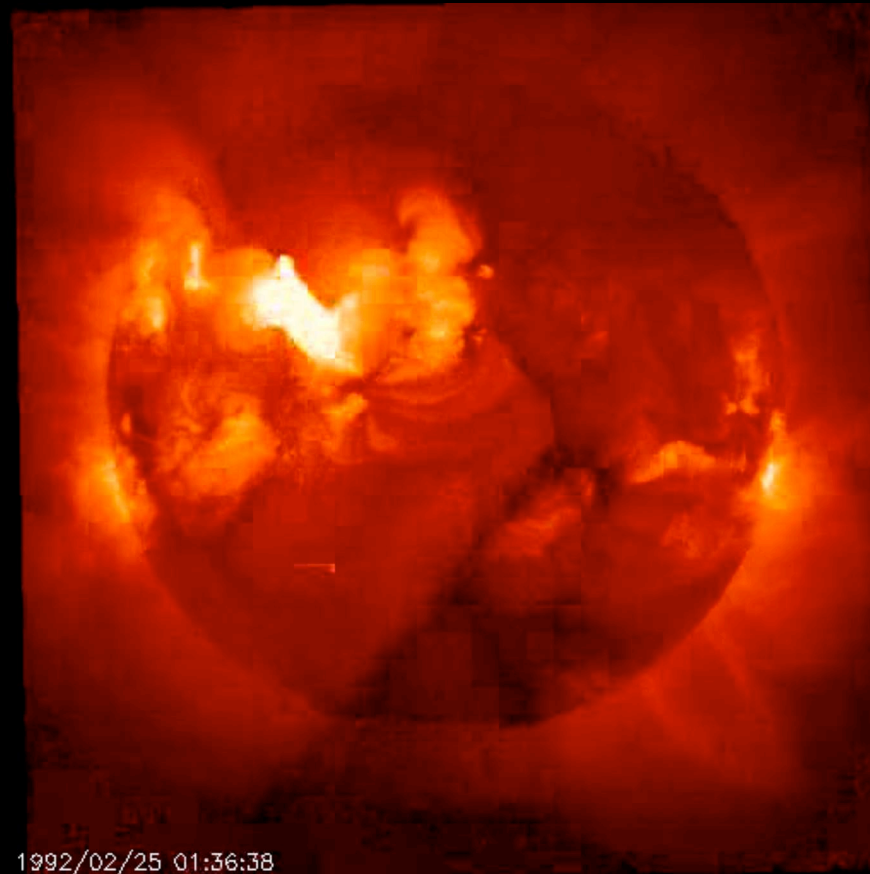
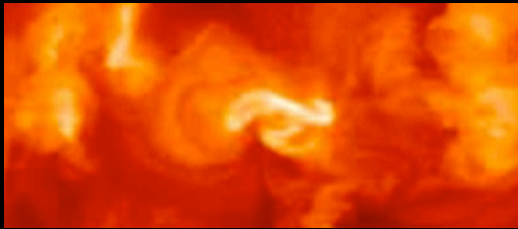
$v_{x1} = 0, v_{x2} = 0$

Magnetic Reconnection

— fast magnetic energy-releasing process —

Solar flare...

explosive phenomenon which
**rapidly releases free magnetic
energy** stored in the corona



Observed by Yohkoh

Evolution of magnetic field

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}$$

evolution caused by the rotation of electric field