## Three components of $E^{MHD} = E_{conv} + E_{res} + E_{Hall}$ All the components are associated with time variation in B

MHD electric field is generated by flow and current, while not by charge (due to local charge neutrality on MHD scale).

 $m{E}_{conv} \equiv - m{u}^{MHD} imes m{B}$  ... related to average flow transverse to magnetic field; caused by gyration of protons and electrons around the magnetic field B\_-component

Convective electric field produces transport of B at the flow velocity u (froze-in). => charged particle gyrates around the same magnetic field line, or magnetic flux inside a gyration orbit is an invariant.

It causes energy conversion between magnetic energy and flow energy  $(E_{conv} \cdot j^{MHD} = (j^{MHD} \times B) \cdot u^{MHD} = 0$ ; reversible process).

$$m{E}_{res} \equiv \eta \; m{j}^{MHD}$$

... related to electric current; caused by collective collisions between protons and electrons (due to charge imbalance via collective collisions, which is different from instantaneous Coulombic electric field arising during a collisional event between a proton and electron)

Resistive electric field produces diffusion of B (non-frozen-in). => charged particles detach from magnetic field line. It causes energy conversion from magnetic energy to thermal energy ( $E_{res} \cdot j^{MHD} > 0$ ; irreversible process => Ohmic heating). For astronomical & space plasmas,  $Rm = \frac{v L}{\eta_{digr}} \sim \frac{E_{come}}{E_{res}}$  is normally quite large, so  $E_{resis}$  is only effective in a local region where  $R_m$  becomes small (e.g. current sheet).



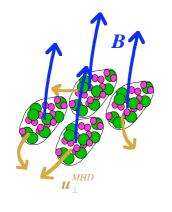
$$m{E}_{Hall} \equiv rac{1}{e \ n} \ m{j}^{MHD} imes m{B}$$

 $E_{Hall} \equiv \frac{1}{e n} j^{MHD} \times B$  ... related to electric current; caused by decoupling between protons and gyrating **electrons** in  $B_1$  – *plane* 

Hall electric field produces slipping of B (different from diffusion). => electrons and magnetic field slip through protons. It does not convert magnetic energy  $(E_{Holl} * j^{MHD} = 0)$ . It is effective in a local region where B is weak so that protons and electrons are decoupled in  $B_1$ -plane (e.g. central part of a current sheet).

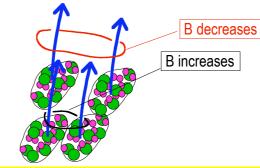
$$\frac{E_{Hall}}{E_{conv}} \sim \frac{t_B}{t_{MHD}} \quad t_B \sim \omega_B^{-1} = \frac{M}{e B}, t_{MHD} \sim \frac{L}{v_A} \quad \left(v_A = \frac{B}{\sqrt{\mu_0 \, \rho}}\right)$$



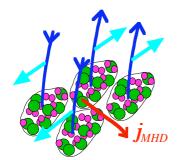


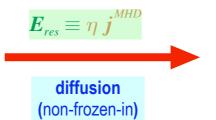


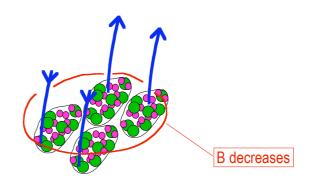
convection (frozen-in to MHD fluid element)



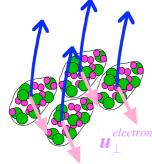
collisions between protons and electrons are less effective

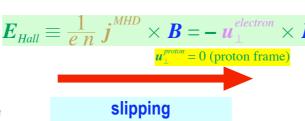




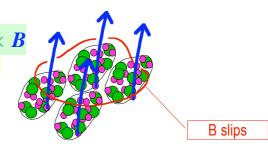


collisions between protons and electrons are effective





slipping (frozen-in to electron fluid element)



collisions between protons and electrons are less effective