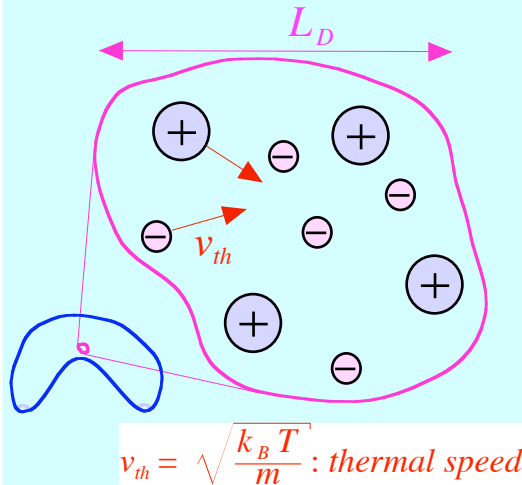


## Typical scales in both approaches...

**Kinetic approach...** particle is a fundamental object (its internal structure is not considered)



**Typical scales:**

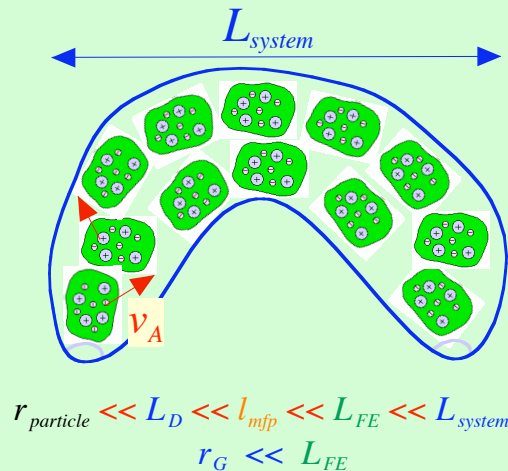
Length...  $L_D$  (Debye length)

e.g., 1 cm for a solar coronal plasma

Time...  $L_D / v_{th} \sim 1 / \nu_p$  (plasma frequency)

e.g.,  $5 \times 10^{-9}$  s for a solar coronal plasma

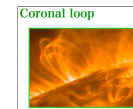
**Fluid approach...** fluid element is a fundamental object (its internal structure is not considered)



**Typical scales:**

Length...  $L_{system}$  (System size)

e.g., 100,000 km  $\sim 10^8$  m for a coronal loop



Time...  $L_{system} / v_A$  ( $v_A = \frac{B}{\sqrt{\mu_0 n m}}$ : Alfvén speed)

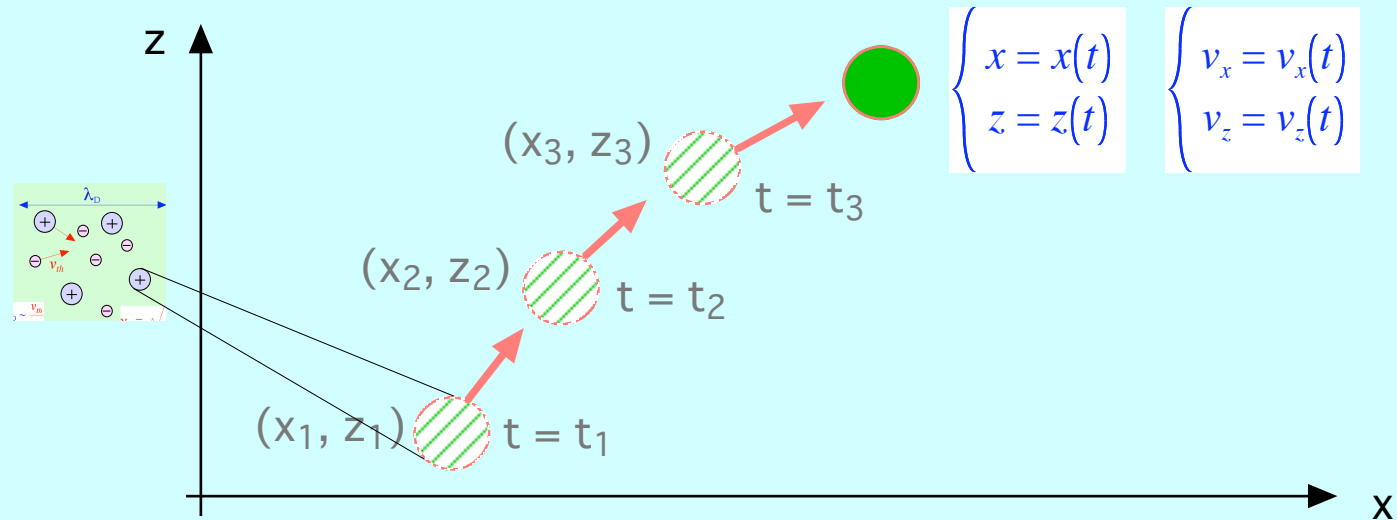
e.g., 100 s for a coronal loop ( $v_A \sim 1000$  km/s)

## Summary of characteristic scales of plasmas

	E-related scale			B-related scale	
Length	mean free path ( $l_{mfp}$ )	Debye length ( $L_D$ )	effective radius ( $r_{eff}$ )	gyration radius ( $r_G$ )	mean interval ( $l$ )
Time	collision time ( $t_c = \frac{l_{mfp}}{v_T}$ )	oscillation period ( $t_p = \frac{L_D}{v_T}$ )		gyration period ( $t_G = \frac{r_G}{v_T}$ )	
Physical process	collision	oscillation	collision with a large scattering angle	gyration	
Physical effect	thermalization	neutralization	thermalization	pressure & current in $B_{\perp}$ -plane	
CGS unit	$l_{mfp}^{e-e} = \frac{(k_B T_e)^2}{4 \pi n_e e^4} \frac{1}{\ln N_D}$ $L_D = \sqrt{\frac{k_B T}{4 \pi e^2 n_e}}$ $r_{eff} \sim \frac{e^2}{k_B T}$ $r_G = \frac{m c v_T}{e B}$ $l \sim n^{-\frac{1}{3}}$				

**Basic equations in kinetic & fluid approaches**

**Kinetic approach...** Focus on the **position** and **velocity** of a **particle** at **every time**  
**=> Mechanical equation**



In the kinetic approach, we consider the **physical state** of a **particle** represented by its **position** and **velocity**.

The **mechanical equation** is a **differential equation** where **time** is the **only independent variable** (ordinary differential equation, ODE).

$$\begin{cases} m \frac{dv_x}{dt} = F_x(x(t), z(t), v_x(t), v_z(t), t) \\ m \frac{dv_z}{dt} = F_z(x(t), z(t), v_x(t), v_z(t), t) \end{cases} \quad \begin{cases} \frac{dx}{dt} = v_x(t) \\ \frac{dz}{dt} = v_z(t) \end{cases}$$